



# Computing Network Configuration Merit

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# Introduction

- Gravitational wave astronomy will require the use of co-operative networks of multiple interferometers.
- We naturally wish to determine not only how to optimally analyze data from a given network, but also how to optimally configure the system as a whole.
- We present a formalism enabling the comparison of the quality of given networks under certain criteria.
- This is achieved by the definition of physically significant *figures of merit* for any particular configuration of a global network and analysis system.
- We give particular consideration to the role the proposed Australian-International Gravitational wave Observatory (AIGO) will play in a global network.



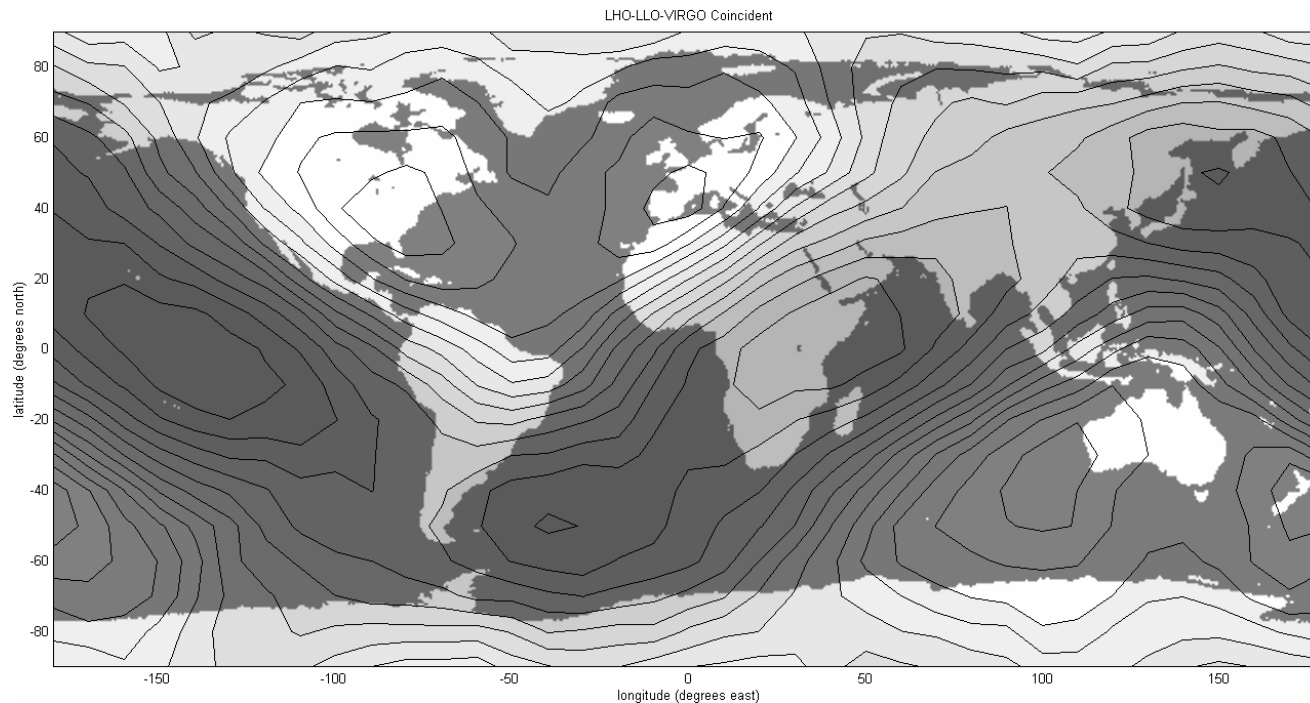
# Overview

- Figure of merit formalism.
- Binary inspiral detection rate as a figure of merit.
  - Maximum likelihood formulation of network analysis algorithms.
    - Approximations and assumptions for computational amenability.
  - Seeing distance, volume and event rate.
  - Generally applicable figure of merit.
  - Application to optimal augmentation of an existing network.
    - Geographical maps of augmentation site merit.
      - Determination of optimal sites, including AIGO.
    - Interpretation of quantitative and qualitative behavior of analyses.
- Future directions and figures of merit under development.
  - Detectable fraction of galactic pulsars as a figure of merit, with Finn.

# Recent Results

■ 4<sup>th</sup> Edoardo Amaldi Conference

■ CQG 19 (2002) 1465-1470





# Figures of Merit

- Compare the quality of networks  $\alpha$  and  $\beta$  by assigning them merits  $f(\alpha)$  and  $f(\beta)$ .
  - $\alpha$  ‘better than’  $\beta$  if and only if  $f(\alpha) > f(\beta)$ .
- A *figure of merit* is a function  $f$  assigning a real value to each member of a set of networks  $\mathcal{N}$ .
  - $f : \mathcal{N} \rightarrow \mathbb{R}$



# Binary Inspiral Detection Rate

- A figure of merit that is strictly increasing with the rate of detections produced, by a population of binary inspiral events, in any network of detectors, using a particular analysis strategy.
  - It is important that the figure of merit also be amenable to computation.

# Binary Inspiral Signal

- The response of a single ideal detector to a binary inspiral is:

- $$m(t) = \frac{\alpha(t)}{r} [(1 + \cos^2 i) \cos \beta(t) F_+ + 2 \cos i \sin \beta(t) F_\times]$$

- Relative orientations are encoded by antenna patterns  $F_+$  and  $F_\times$ .  $r$  and  $i$  are source distance and orbital inclination.  $\alpha(t)$  and  $\beta(t)$  depend on other source properties (such  $m_1$  and  $m_2$ ).



# Network Analysis Algorithms

- Coincident and coherent network analyses can be similarly implemented as likelihood tests.
  - L. S. Finn, Phys Rev D **63** 102001 (2001)
  - We implement the simplest forms of each test.
    - The performance of both tests could be improved at the cost of added complexity.
    - Coincident veto strategy will be unimportant as we will be concerned only with false dismissals.





# Maximum Likelihood Tests

- The likelihood  $\Lambda$  that a signal response  $m(t)$  is present in the output  $g(t)$  of a detector is:
  - $\ln \Lambda(g|m) = \langle 2g - m, m \rangle$
- The maximum likelihood  $\Lambda_{\max}$  is the maximum over a set of signal responses  $\mathcal{M}$ :
  - $\Lambda_{\max}(g|\mathcal{M}) = \max_{m \in \mathcal{M}} \Lambda(g|m)$
- Detection occurs when a threshold is exceeded:
  - $\Lambda_{\max} > \Lambda$



# Coincident Algorithm

- A coincident analysis makes a detection when each detector in a network crosses a threshold:
  - $\Lambda_{\max}(g_i|\mathcal{M}) > \Lambda_i$
  - If each detector in the network is identical and hence has same threshold:
    - $\min_i \Lambda_{\max}(g_i|\mathcal{M}) > \Lambda_{\text{coincident}}$
    - Further vetoes may be applied.



# Coherent Test

- A coherent analysis makes a detection when the network as a whole crosses a threshold:
  - $\Lambda_{\max}(\mathbf{g}|\mathcal{M}) > \Lambda_{\text{coherent}}$
  - If the noise is uncorrelated between detectors the network's likelihood is separable:
    - $\max_{m \in \mathcal{M}} \prod_i \Lambda(g_i|m) > \Lambda_{\text{coherent}}$
    - A. Pai *et al*, gr-qc/0110041



# False Dismissal

- We are concerned with the false dismissal of events from the population.
  - The selection of thresholds reduces false alarms to an acceptable (insignificant) rate.
- We are then interested in the case when the detector output consists of a signal and additive noise.
  - $g(t) = m(t) + n(t)$

# Approximation

- When a signal  $m$  is present and detected:
  - $\Lambda_{\max}(m + n|\mathcal{M}) > \Lambda_0$
- Approximate mean using the real signal:
  - $$\begin{aligned}\overline{\ln \Lambda_{\max}(m + n|\mathcal{M})} &\approx \overline{\ln \Lambda(m + n|m)} \\ &= \overline{\langle m + 2n, m \rangle} \\ &= \langle m, m \rangle\end{aligned}$$

# Approximated Tests

- For a coincident analysis:

$$\square \min_i \overline{\ln \Lambda_{\max}(m_i + n_i | \mathcal{M})} \approx \min_i \langle m_i, m_i \rangle$$
$$> \ln \Lambda_{\text{coincident}}$$

- For a coherent analysis:

$$\square \max_{m \in \mathcal{M}} \sum_i \overline{\ln \Lambda(m_i + n_i | m_i)} \approx \sum_i \langle m_i, m_i \rangle$$
$$> \ln \Lambda_{\text{coherent}}$$

# Seeing Distance and Volume

- Consider a particular source from the population, fixed except for distance  $r$ .

- Determine the maximum detectable distance:

- $\min_j \langle m_j, m_j \rangle = \ln \Lambda_{\text{coincident}}$

- $$r^2 = \frac{\langle \alpha \cos \beta, \alpha \cos \beta \rangle}{\ln \Lambda_{\text{coincident}}} \min_j [(1 + \cos^2 i)(F_+^2)_j + 4 \cos^2 i (F_\times^2)_j]$$

- $\sum_i \langle m_i, m_i \rangle = \ln \Lambda_{\text{coherent}}$

- $$r^2 = \frac{\langle \alpha \cos \beta, \alpha \cos \beta \rangle}{\ln \Lambda_{\text{coherent}}} \sum_j [(1 + \cos^2 i)(F_+^2)_j + 4 \cos^2 i (F_\times^2)_j]$$

- Average over source orientations and integrate over the sky for detectable volume, and thus event rate.

# Simplifications

- The leading constant does not depend on the geographical configuration of the detectors, only on the source waveform  $(\alpha, \beta)$  and the noise statistics  $(\Lambda)$ .

- $$V = \frac{\langle \alpha \cos \beta, \alpha \cos \beta \rangle}{\ln \Lambda_{\text{threshold}}} \int \dots d\Omega$$

- We make a general figure of merit by eliminating the leading constant.



# Figure of Merit

- The final form of the figures of merit for:

- Coincident:

$$\int \left\{ \min_j [(1 + \cos^2 i)(F_+^2)_j + 4 \cos^2 i (F_\times^2)_j] \right\}^{\frac{3}{2}} d\Omega$$

- Coherent:

$$\int \left\{ \sum_j [(1 + \cos^2 i)(F_+^2)_j + 4 \cos^2 i (F_\times^2)_j] \right\}^{\frac{3}{2}} d\Omega$$

- Not the detection rate, but linearly proportional to it.

- Independent of non-geometric source and detector properties, and so valid for any inspiral parameters and any detector noise PSD.



# Network Configurations

- Define the set  $\mathcal{N}_n$  of networks of  $n$  detectors each identical up to location and orientation.
  - $\mathcal{N}_n = \{[(\theta_1, \phi_1, \psi_1), \dots, (\theta_n, \phi_n, \psi_n)]\}$
- Each detector  $i$  is completely defined by its:
  - Latitude  $\theta_i$ .
  - Longitude  $\phi_i$ .
  - Orientation  $\psi_i$ .



# Configurations of Interest

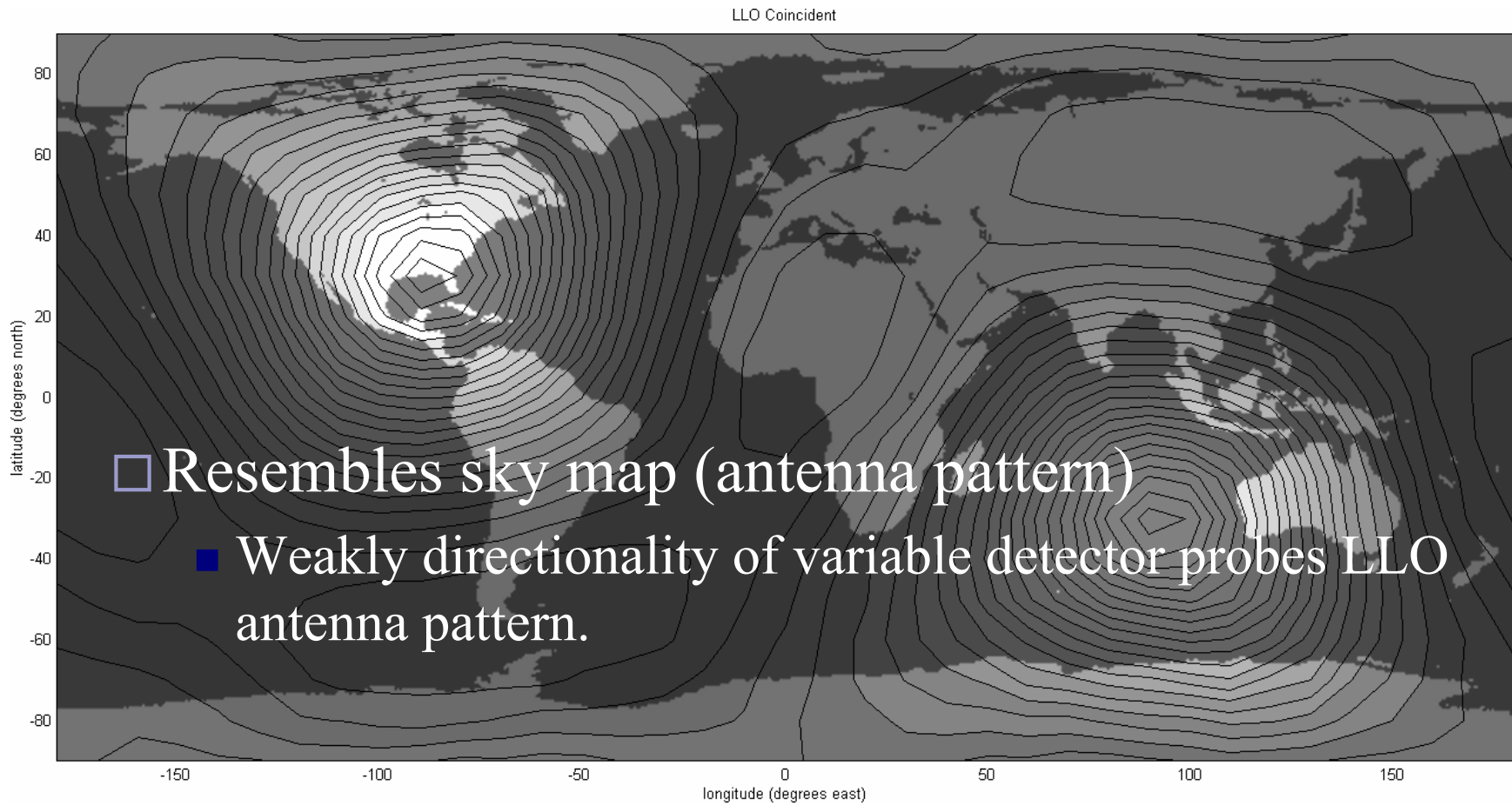
- Consider how to best augment an existing network of  $(n - 1)$  identical detectors with an  $n$ th identical detector for inspiral detection.
  - Fix  $(n - 1)$  detectors, vary  $n$ th detector.
    - Forms a 3 dimensional subset of  $\mathcal{N}_n$ .
  - Note that merit is insensitive to orientation  $\psi_n$  which we fix at an arbitrary value.
    - Forms a 2 dimensional subset of  $\mathcal{N}_n$ .



# Maps of Augmentation Site Merit

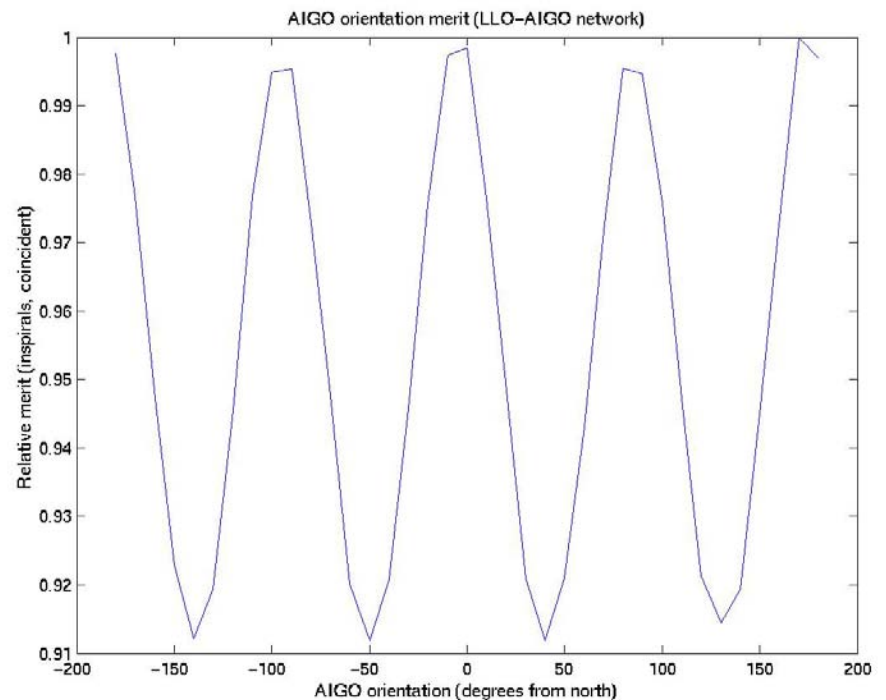
- We then compute the figures of merit for points in the two dimensional subset parameterized by the latitude and longitude of the  $n$ th detector.
  - This set has the straightforward interpretation as a geographical map of the merit of a particular site for augmenting an existing network.
- Integrals evaluated by Monte Carlo methods using the APAC National Facility supercomputer.

# Coincident: LLO

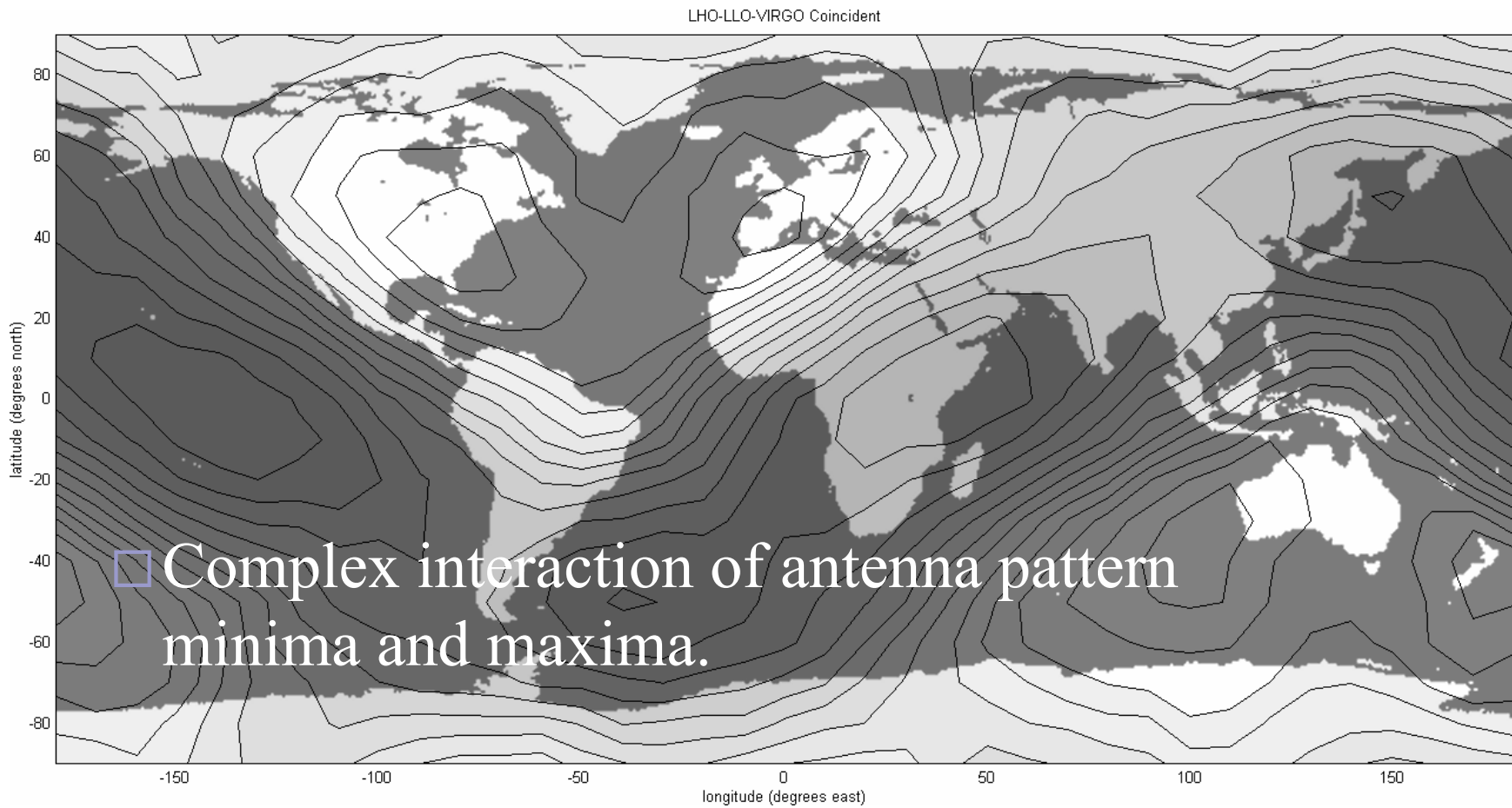


# Weak Orientation Dependence

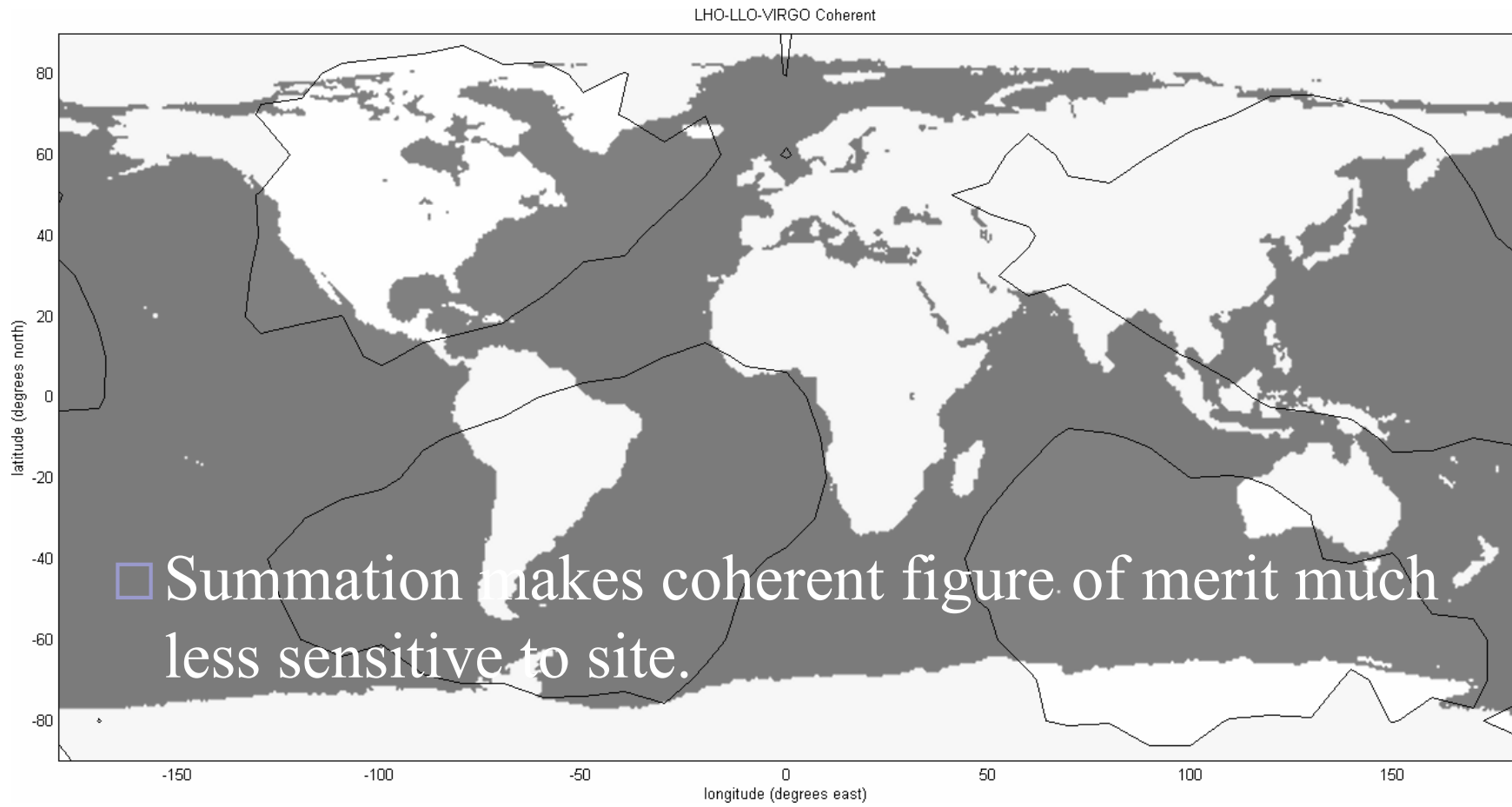
- Rotating a detector at the AIGO site varies the merit in the preceding case by  $< 10\%$ , compared with a  $60\%$  variation from varying the latitude and longitude.
- Other circumstances produce even smaller variations.



# Coincident: LLO + LHO-4k + VIRGO



# Coherent: LLO + LHO-4K + VIRGO







# Interpretation

## ■ Results:

- Coincident analysis sensitive to detector sites.
  - AIGO well placed.
  - N. Arnaud *et al*, gr-qc/0107081
- Coherent analysis insensitive to detector sites.
- Realistic networks will be closer to optimal for a coherent analysis.

## ■ Limitations:

- We can only compare networks with the same number of detectors and the same analysis method.
  - Cannot say if coherent or coincident would produce a higher absolute rate.
  - Cannot say how beneficial adding a detector is.



# Future Directions

- We are producing new figures of merit to answer other questions.
  - Improving likelihood approximation to extend binary inspiral detection rate figure of merit applicability.
  - Angular resolution of a network as a figure of merit.



# Summary

- We have developed a general formalism for comparing networks and analysis methods.
- Geographical (mis)configuration is a significant factor affecting the detection of binary inspirals for a coincident network analysis, but much less so for a coherent network analysis.
- We are currently working on implementing new and improved figures of merit.